



Improving Pulsar Observations by Training a Machine Learning Model on the EPTA Dataset

Jacob Cardinal Tremblay¹, Jędrzej Jawor²
¹Max Planck Institute for Gravitational Physics (Albert Einstein Institute)
²Max Planck Institute for Radio Astronomy



EPTA Data and Pipeline

- Up to 25 years of observations (EPTA Collaboration et al. 2023)
 - High-quality folded data from the Effelsberg 100-meter radio telescope
- New pipeline for Radio Frequency Interference (RFI) cleaning
 - RCVR → CLFD → MEERSurg → Manual Cleaning
- Statistical algorithms need variable thresholding
 - Manual oversight required
 - Non-standardized way of cleaning
- Ultra-Broadband Receiver (UBB)
 - Future cleaning will be much more intensive
- Improve Time of Arrival (TOA) uncertainties by:
 - Removing more RFI
 - Eliminating less “clean” data



Radio Frequency Interference

- Radio Frequency Interference (RFI) = Unwanted human-made signals
- Large frequency bandwidth = A lot of RFI!
- RFI can reduce quality or even completely corrupt data!
- Can be narrow-band or wide-band

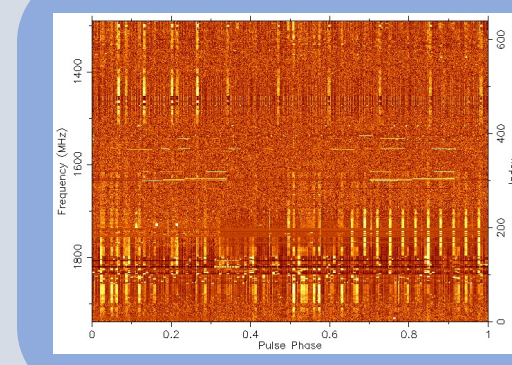


Image Credit: Selman Icons

Machine Learning

- Train a model on EPTA dataset to identify RFI
 - Random Forest/Gradient Boosting are ideal this case
 - E.g. A. Berthereau et al. 2023
- Obtain weights from previously cleaned observations
- Calculate statistics on raw data
 - E.g. median, mean, skewness, kurtosis
- Use statistics as features on which to train
 - Do not train on the actual raw data
- 80% of the data for training, 20% for testing

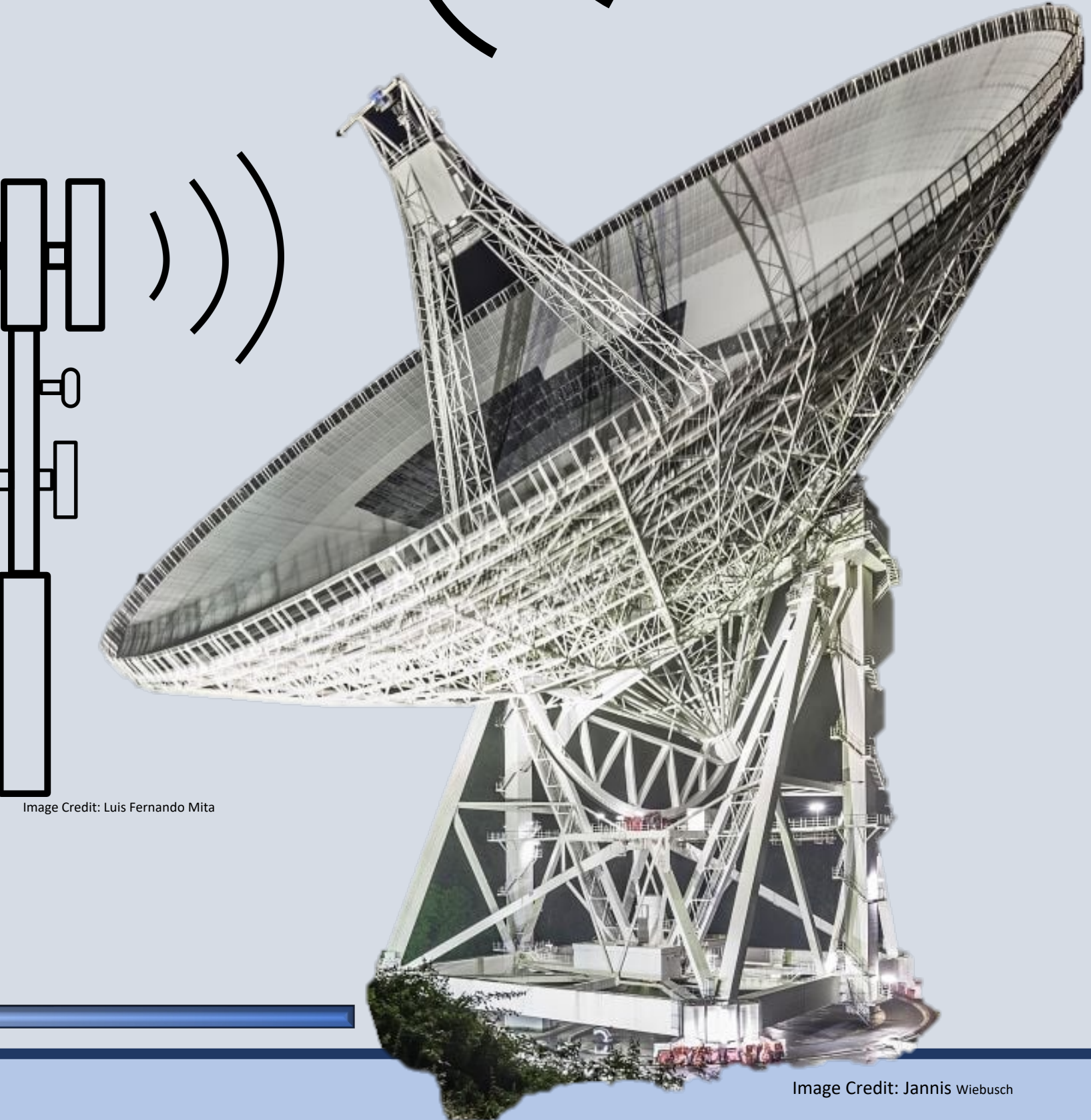
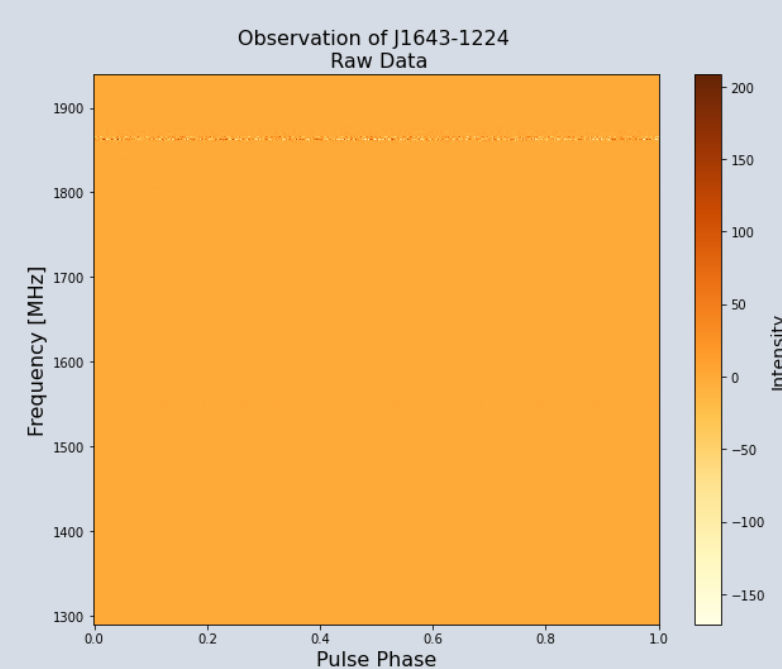
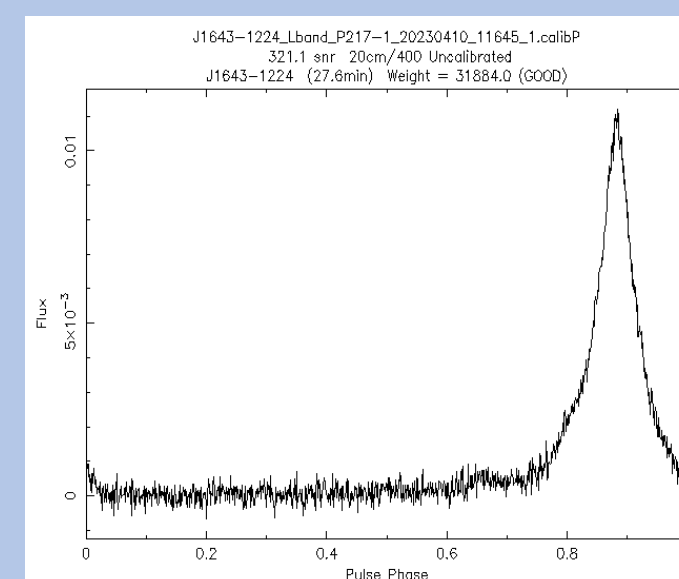


Image Credit: Jannis Wiebusch

Testing

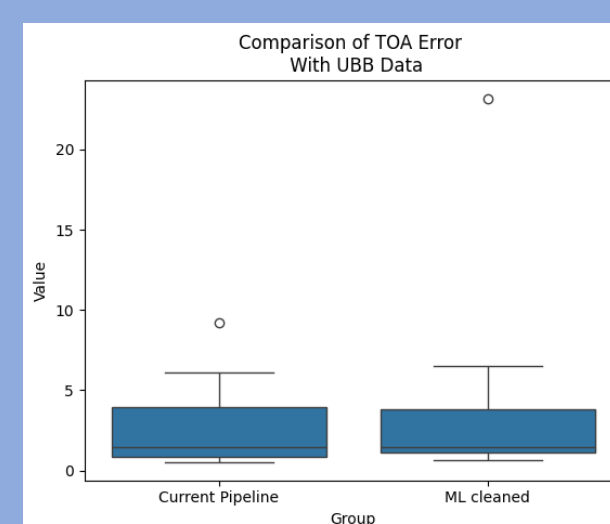
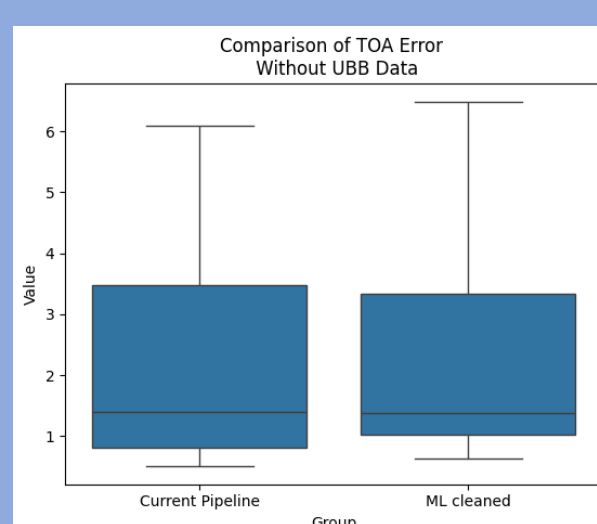
- Initial tests on pulsar J1643-1224
 - Very bright pulsar
 - “Single-peak” profile
- 74 EDD observations → ~ 6 million individual frequency channels
- Dataset is relatively balanced
 - 44% RFI, 56% no RFI



- Tested ~16,000 combinations of features to arrive at best features based on accuracy and F1-Score
 - frequency, median, mean, skewness, max

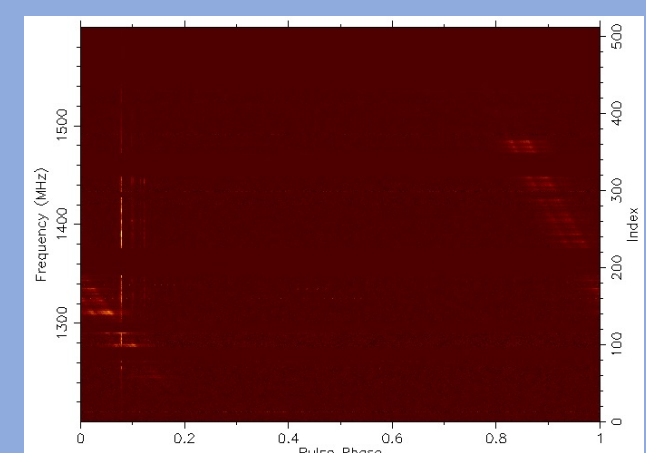
Current Results

- Calculated metrics:
 - Accuracy: 94.06%
 - Precision: 90.62%
 - Recall: 99.17%
 - F1-Score: 94.70%
- TOAs calculated from cleaning
 - Uncertainties in TOAs are compared
 - ML cleaning is slightly more consistent
- However, performs worse on UBB data



Future Work

- Explore more machine learning models?
- Hyperparameter tuning
- Train on the off-pulse
- Train on multiple backends
- Train on multiple pulsars
- Train and/or test model on single pulse data
- Include wideband RFI elimination



References

EPTA collaboration, The second data release from the European Pulsar Timing Array I. The dataset and timing analysis, Astron. Astrophys. 678 (2023) A48 [2306.16224].
Anaïs Berthereau. Méthodes avancées en chronométrie des pulsars : réjection d’interférences radio par apprentissage profond et étude chronométrique du système binaire relativiste J1528-3146. Astrophysique [astro-ph]. Université d’Orléans, 2023. Français. finNT : 2023ORLE1043ff. ftitel-04509283
Morello, V., Barr, E. D., Cooper, S., et al. 2019, MNRAS, 483, 3673
MeerGuard: <https://github.com/danielreardon/MeerGuard>
CoastGuard3: <https://gitlab.mpcdf.mpg.de/jawor/CG-3>

Acknowledgements

Special thanks to: Marie Dumaz, Rutger van Haasteren, Boris Goncharov, Karsten Beismann, David Champion and Michael Kramer for the various roles they played in allowing this work to become a reality.